

4E : The Quantum Universe



Lecture 26, May 17

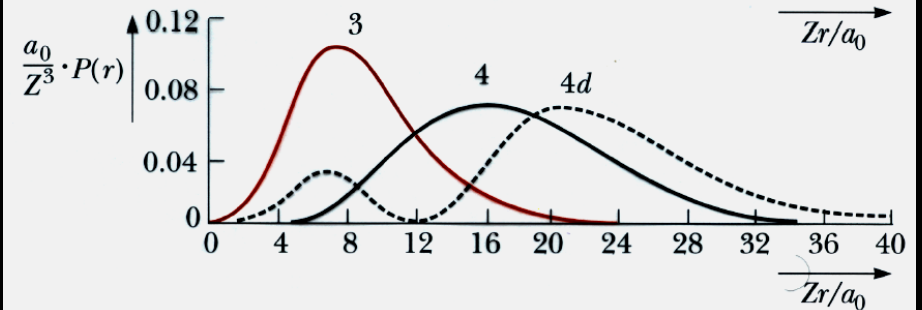
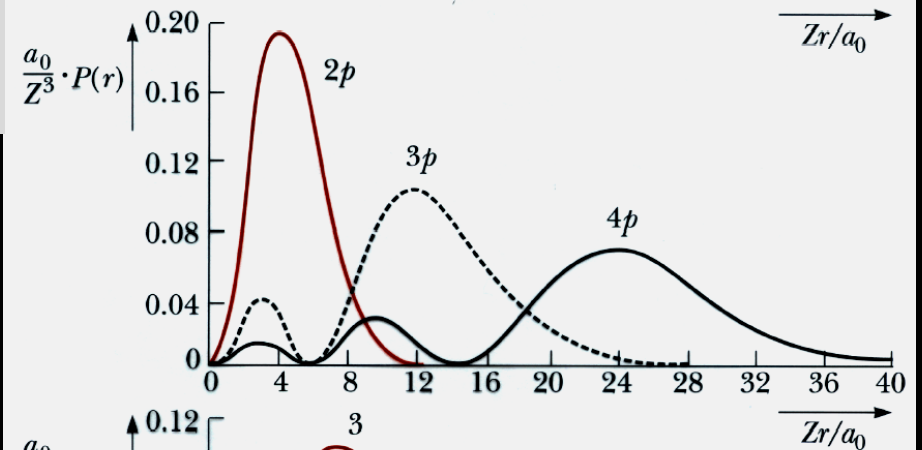
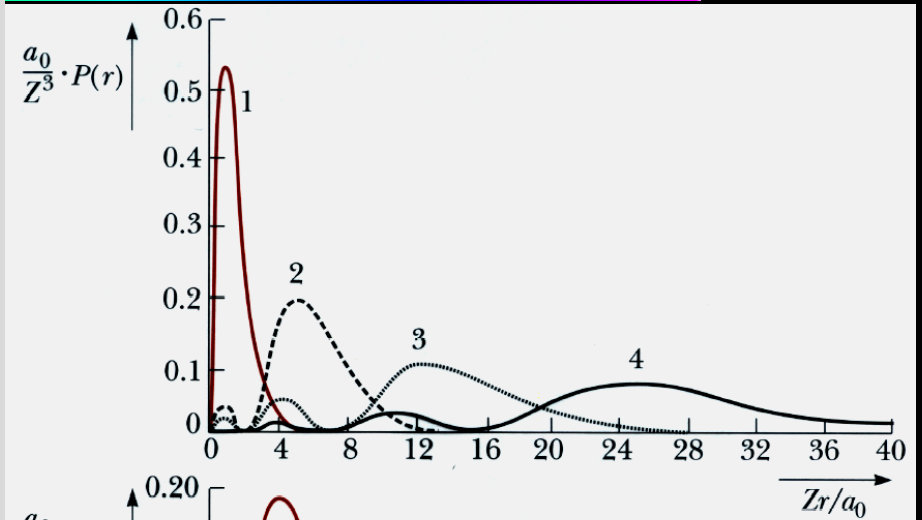
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Radial Probability Distribution $P(r) = r^2 R(r)$

TABLE 7-2 Radial functions for hydrogen

$n = 1$	$l = 0$	$R_{10} = \frac{2}{\sqrt{a_0^3}} e^{-r/a_0}$
$n = 2$	$l = 0$	$R_{20} = \frac{1}{\sqrt{2a_0^3}} \left(1 - \frac{r}{2a_0}\right) e^{-r/2a_0}$
	$l = 1$	$R_{21} = \frac{1}{2\sqrt{6a_0^3}} \frac{r}{a_0} e^{-r/2a_0}$
$n = 3$	$l = 0$	$R_{30} = \frac{2}{3\sqrt{3a_0^3}} \left(1 - \frac{2r}{3a_0} + \frac{2r^2}{27a_0^2}\right) e^{-r/3a_0}$
	$l = 1$	$R_{31} = \frac{8}{27\sqrt{6a_0^3}} \frac{r}{a_0} \left(1 - \frac{r}{6a_0}\right) e^{-r/3a_0}$
	$l = 2$	$R_{32} = \frac{4}{8\sqrt{30a_0^3}} \frac{r^2}{a_0^2} e^{-r/3a_0}$



Because $P(r) = r^2 R(r)$; No matter what $R(r)$ is for some n ,
The prob. Of finding electron inside the nucleus = 0 !!

Normalized Spherical Harmonics & Structure in H Atom

TABLE 7-1 Spherical harmonics

$l = 0$	$m = 0$	$Y_{00} = \sqrt{\frac{1}{4\pi}}$
$l = 1$	$m = 1$	$Y_{11} = -\sqrt{\frac{3}{8\pi}} \sin \theta e^{i\phi}$
	$m = 0$	$Y_{10} = \sqrt{\frac{3}{4\pi}} \cos \theta$
	$m = -1$	$Y_{1-1} = \sqrt{\frac{3}{8\pi}} \sin \theta e^{-i\phi}$
$l = 2$	$m = 2$	$Y_{22} = \sqrt{\frac{15}{32\pi}} \sin^2 \theta e^{2i\phi}$
	$m = 1$	$Y_{21} = -\sqrt{\frac{15}{8\pi}} \sin \theta \cos \theta e^{i\phi}$
	$m = 0$	$Y_{20} = \sqrt{\frac{5}{16\pi}} (3 \cos^2 \theta - 1)$
	$m = -1$	$Y_{2-1} = \sqrt{\frac{15}{8\pi}} \sin \theta \cos \theta e^{-i\phi}$
	$m = -2$	$Y_{2-2} = \sqrt{\frac{15}{32\pi}} \sin^2 \theta e^{-2i\phi}$

Excited States ($n > 1$) of Hydrogen Atom : Birth of Chemistry !

Features of Wavefunction in θ & ϕ :

Consider $n = 2, l = 0 \Rightarrow \psi_{200}$ = Spherically Symmetric (last slide)

Excited States (3 & each with same E_n) :

$\psi_{211}, \psi_{210}, \psi_{21-1}$ are all **2p** states

$$\psi_{211} = R_{21} Y_1^1 = \left(\frac{1}{\pi}\right) \left(\frac{Z}{a_0}\right)^{3/2} \left(\frac{Z}{8}\right) \left(\frac{r}{a_0}\right) e^{-Zr/a_0} \cdot \sin \theta \cdot e^{i\phi}$$

$$|\psi_{211}|^2 = |\psi_{211}^* \psi_{211}| \propto \sin^2 \theta \quad \text{Max at } \theta = \frac{\pi}{2}, \text{ min at } \theta = 0; \text{ Symm in } \phi$$

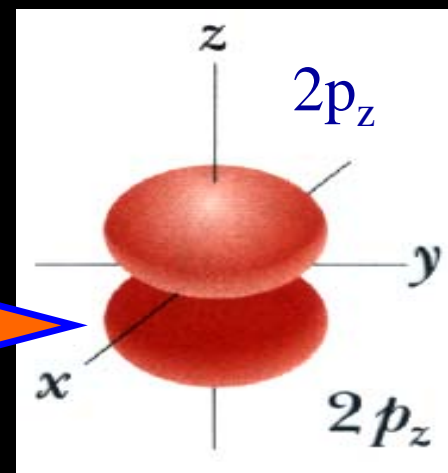
What about ($n=2, \ell=1, m_1 = 0$)

$$\psi_{210} = R_{21}(r) Y_1^0(\theta, \phi);$$

$$Y_1^0(\theta, \phi) \propto \frac{1}{2} \sqrt{\frac{3}{\pi}} \cos \theta;$$

Function is max at $\theta=0$, min at $\theta = \frac{\pi}{2}$

We call this $2p_z$ state because of its extent in z



Excited States ($n > 1$) of Hydrogen Atom : Birth of Chemistry !

Remember Principle of Linear Superposition

for the TISE which is basically a simple differential equation:

$$-\frac{\hbar^2}{2m}\nabla^2\psi + U\psi = E\psi$$

Principle of Linear Superposition \Rightarrow If ψ_1 and ψ_2 are sol. of TISE then a "designer" wavefunction made of linear sum

$\psi' = a\psi_1 + b\psi_2$ is also a sol. of the diff. equation !

To check this, just substitute ψ' in place of ψ
& convince yourself that

$$-\frac{\hbar^2}{2m}\nabla^2\psi' + U\psi' = E\psi'$$

The diversity in Chemistry and Biology **DEPENDS**
on this superposition rule

Designer Wave Functions: Solutions of S. Eq !

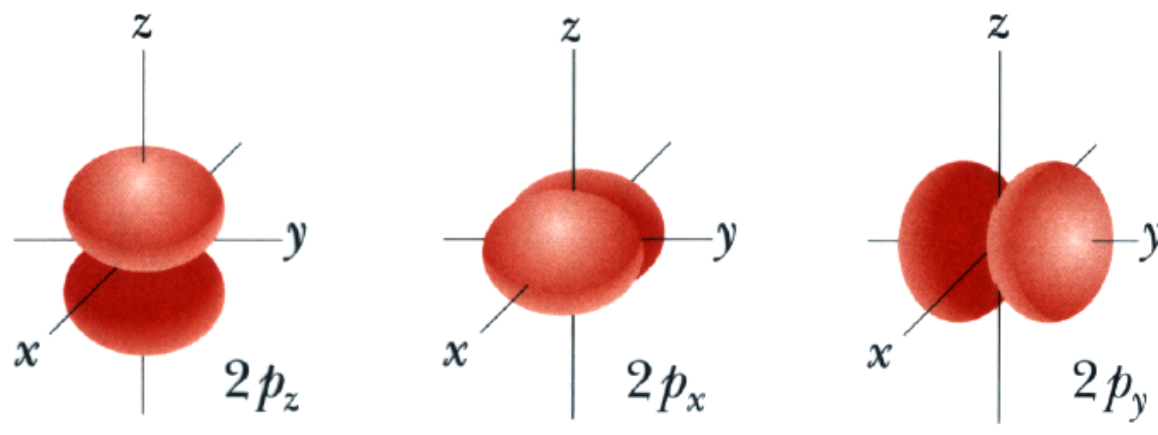
Linear Superposition Principle means allows me to "cook up" wavefunctions

$$\psi_{2p_x} = \frac{1}{\sqrt{2}} [\psi_{211} + \psi_{21-1}] \text{has electron "cloud" oriented along x axis}$$

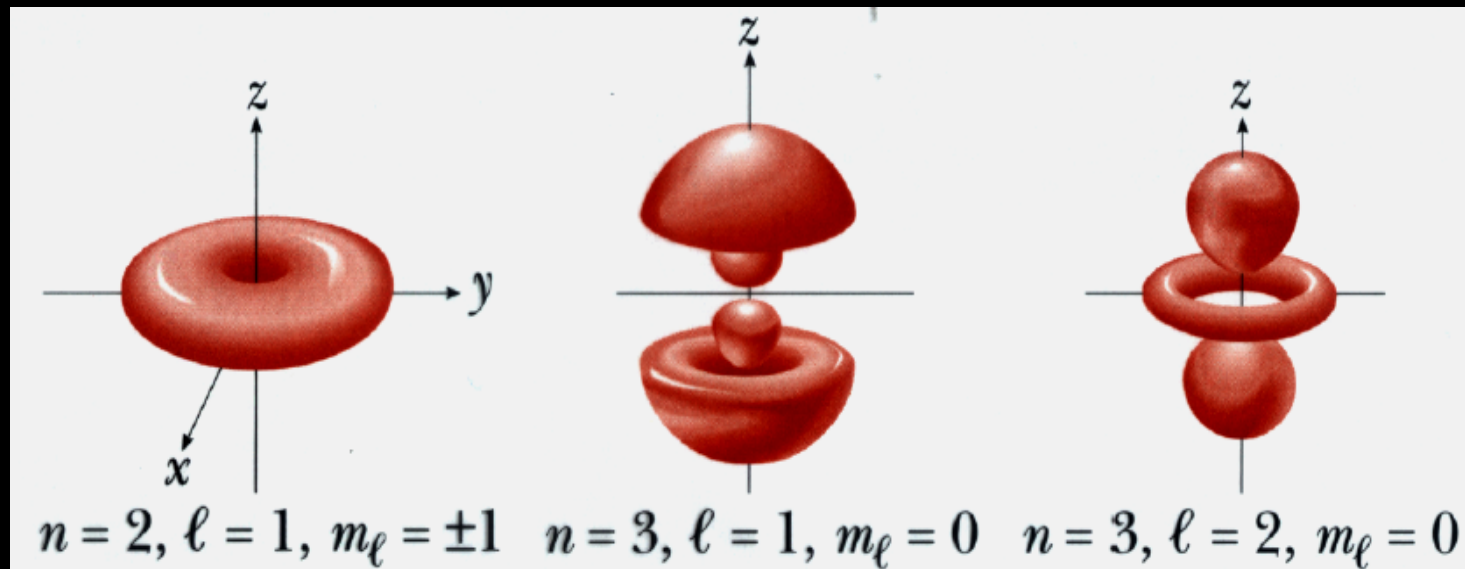
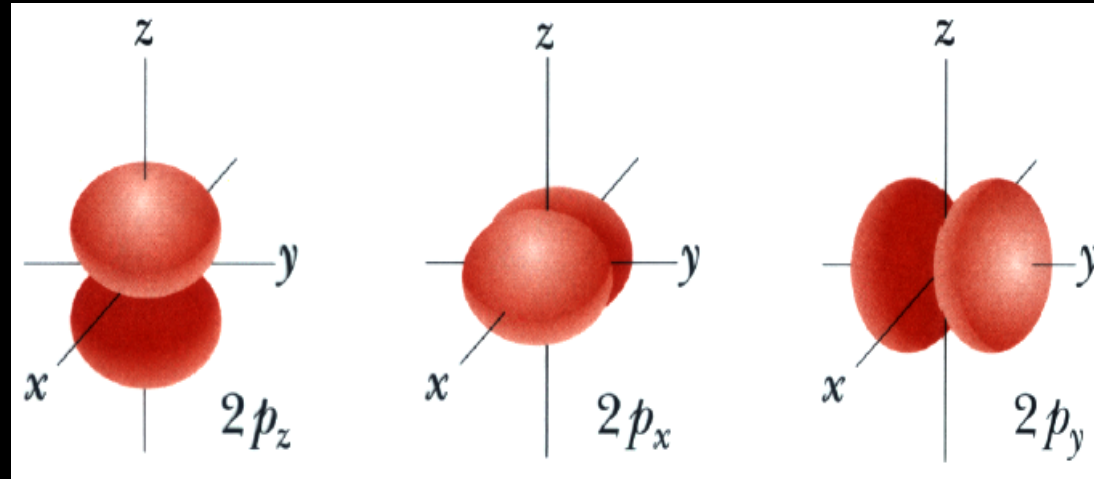
$$\psi_{2p_y} = \frac{1}{\sqrt{2}} [\psi_{211} - \psi_{21-1}] \text{has electron "cloud" oriented along y axis}$$

So from 4 solutions $\psi_{200}, \psi_{210}, \psi_{211}, \psi_{21-1} \rightarrow 2s, 2p_x, 2p_y, 2p_z$

Similarly for n=3 states ...and so on ...can get very complicated structure in θ & ϕwhich I can then mix & match to make electrons "most likely" to be where I want them to be !

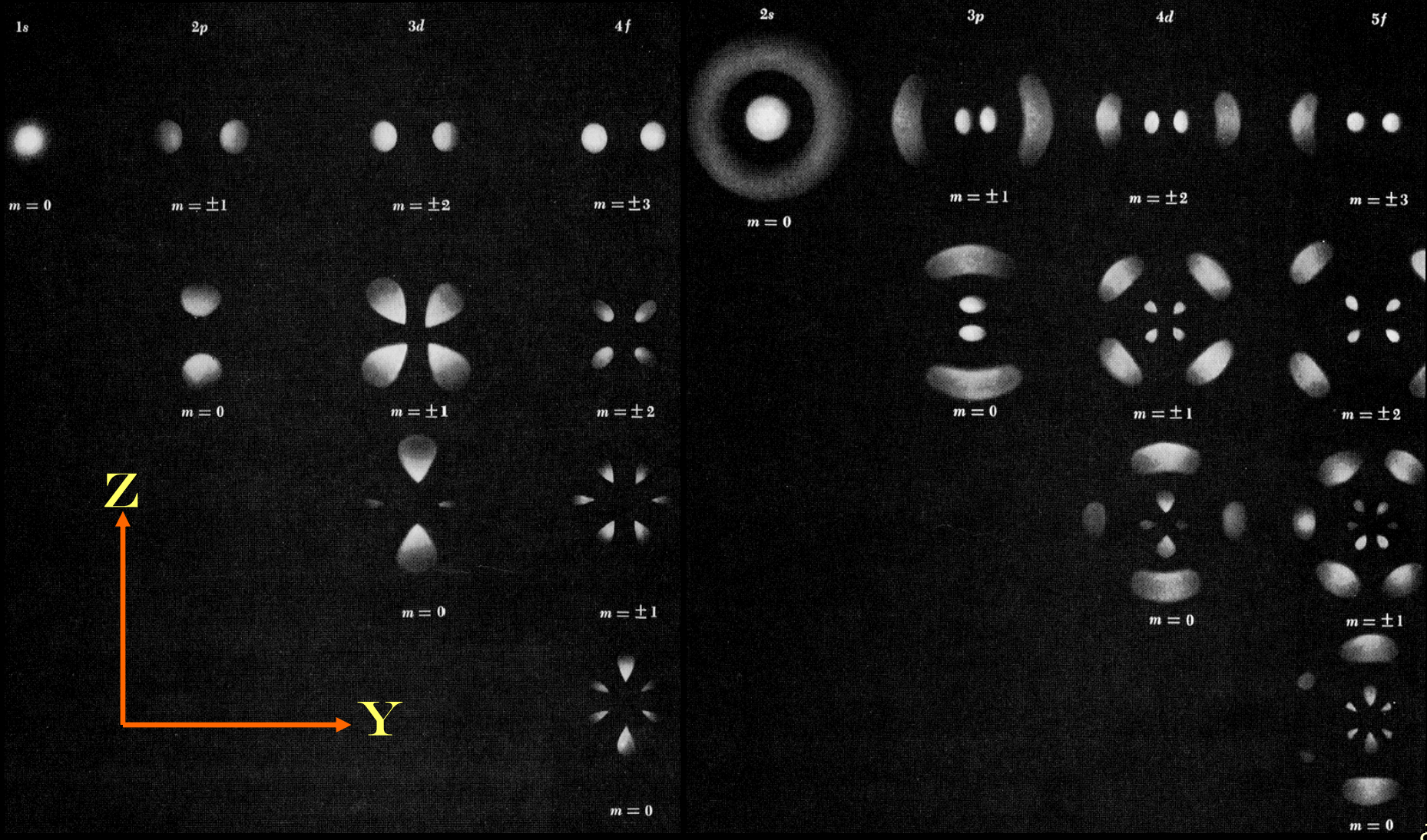


Designer Wave Functions: Solutions of S. Eq !

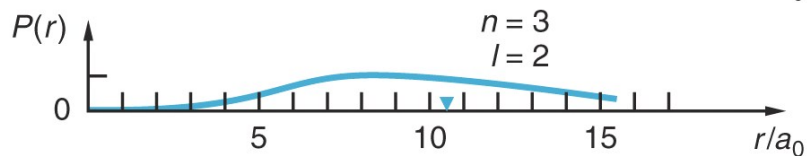
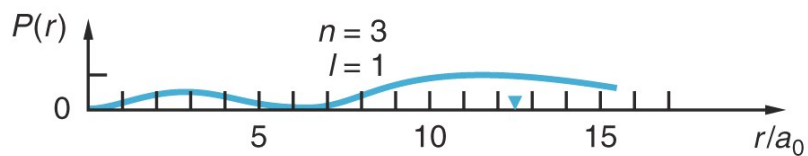
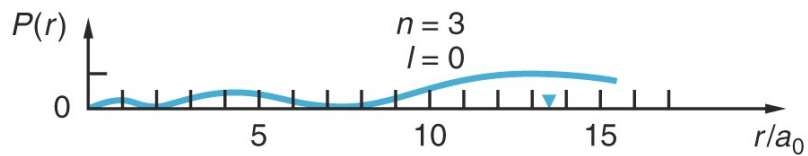
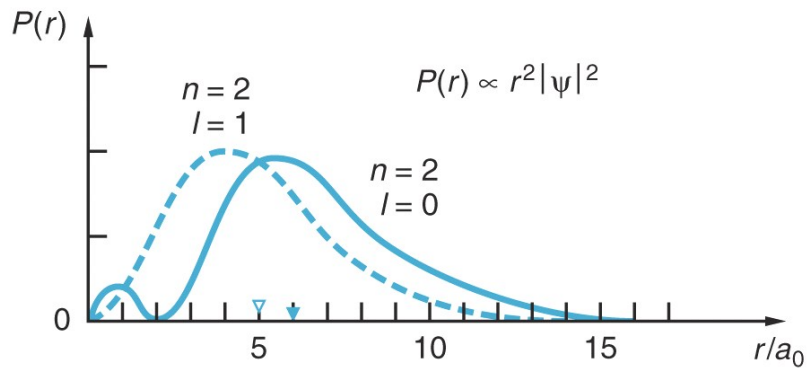


Cross Sectional View of Hydrogen Atom prob. densities in r, θ, ϕ
Birth of Chemistry (Can make Fancy Bonds \rightarrow Overlapping electron "clouds")

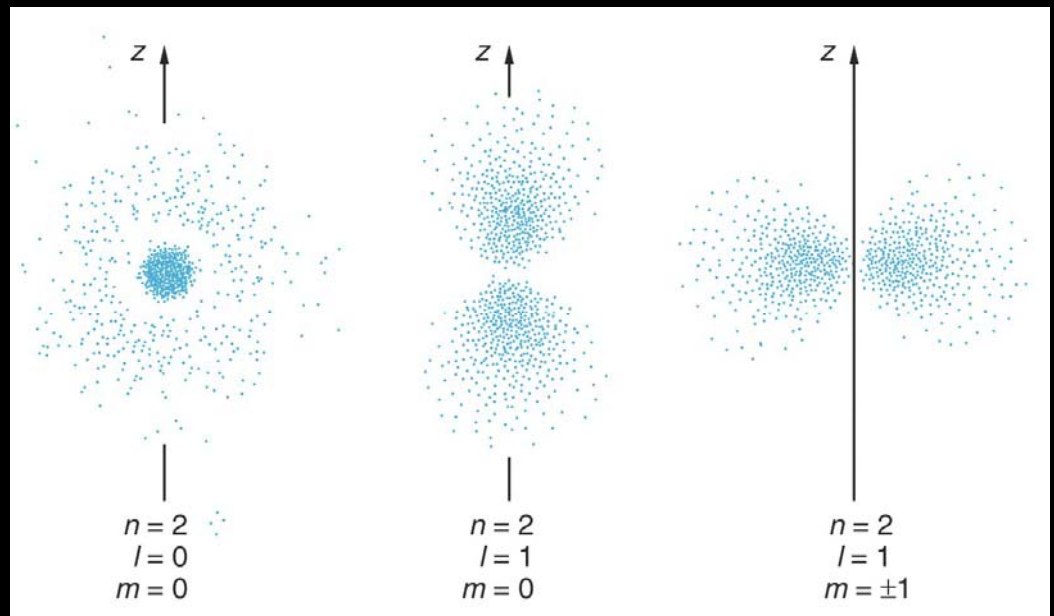
What's the electron "cloud" : Its the Probability Density in r, θ, ϕ space!



More Radial Probabilities $P(r)$ Vs. r/a_0



Net Prob. densities for $n=2$ states



spherically
symmetric

dumbbell

Doughnut
(toroid)

Radiative Transitions: An Impromptu Discussion



- Concepts
 - Write and use the wavefunctions which arise as solutions of the time-dependent Schrodinger equation
 - Use concept of linear superposition of time-dependent wavefunctions
 - Calculate expectation value for $\langle x \rangle$, when is it time-dependent?
- See slides (posted) from Lecture 27 (See also Modern Physics by Beiser, chapter 6, sections 8-9, page 214. Text available from the E-reserve page as pdf files