

# Physics 2D, Winter 2005

## Week 1 Exercise Solutions: Relativity

### 1-2 IN THE REST FRAME:

In an elastic collision, energy and momentum are conserved.

$$p_i = m_1 v_{1i} + m_2 v_{2i} = p_f = m_1 v_{1f} + m_2 v_{2f}$$

or  $m_1(v_{1i} - v_{1f}) = -m_2(v_{2i} - v_{2f})$ . The energy equation is

$$E_i = E_f = \frac{1}{2} m_1 v_{1i}^2 + \frac{1}{2} m_2 v_{2i}^2 = \frac{1}{2} m_1 v_{1f}^2 + \frac{1}{2} m_2 v_{2f}^2$$

or  $m_1(v_{1i} - v_{1f})(v_{1i} + v_{1f}) = -m_2(v_{2i} - v_{2f})(v_{2i} + v_{2f})$ . Substituting the momentum equation into the energy equation yields a very simple and general result (true even for three-dimensional collision if the velocities are replaced as vectors)  $(v_{1i} + v_{1f}) = (v_{2i} + v_{2f})$  or as Newton put it originally, the final relative velocity is opposite to the initial relative velocity:

$$(v_{1i} - v_{2i}) = -(v_{1f} - v_{2f})$$

for elastic collision (and a fraction of the initial for general collisions). Now, putting in the numerical values, the momentum equation, and this relative velocity equation gives:

$m_1 v_{1f} + m_2 v_{2f} = 0.3v_{1f} + 0.2v_{2f} = 0.9 \text{ kg} \cdot \text{m/s}$ , and  $(5 - (-3)) = -(v_{1f} - v_{2f})$ . Solving the two equations, two unknowns, we find  $v_{1f} = -1.4 \text{ m/s}$  and  $v_{2f} = +6.6 \text{ m/s}$ .

### IN THE MOVING FRAME:

The Galilean velocity transformations hold.

$$v'_{1i} - v' = 20 \text{ m/s} - 10 \text{ m/s} = 10 \text{ m/s}$$

$$v'_{2i} = v_{2i} - v' = 0 \text{ m/s} - 10 \text{ m/s} = -10 \text{ m/s}$$

$$p'_i = m_1 v'_{1i} + m_2 v'_{2i} = (2000 \text{ kg})(10 \text{ m/s}) - (1500 \text{ kg})(10 \text{ m/s}) = 5 \times 10^3 \text{ kg} \cdot \text{m/s}$$

$$p'_f = (2000 \text{ kg} + 1500 \text{ kg})v'_f = (3500 \text{ kg})(v_f - 10 \text{ m/s}), \text{ and because } v_f = 11.4 \text{ m/s},$$

$$p'_f = 5 \times 10^3 \text{ kg} \cdot \text{m/s}$$

### 1-3 IN THE REST FRAME:

In an elastic collision energy and momentum are conserved.

$$p_i = m_1 v_{1i} + m_2 v_{2i} = (0.3 \text{ kg})(5 \text{ m/s}) + (0.2 \text{ kg})(-3 \text{ m/s}) = 0.9 \text{ kg} \cdot \text{m/s}$$

$$p_f = m_1 v_{1f} + m_2 v_{2f}$$

This equation has two unknowns, therefore, apply the conservation of kinetic energy

$E_i = E_f = \frac{1}{2} m_1 v_{1i}^2 + \frac{1}{2} m_2 v_{2i}^2 = \frac{1}{2} m_1 v_{1f}^2 + \frac{1}{2} m_2 v_{2f}^2$  and conservation of momentum one finds that  $v_{1f} = -1.31 \text{ m/s}$  and  $v_{2f} = 6.47 \text{ m/s}$  or  $v_{1f} = -1.56 \text{ m/s}$  and  $v_{2f} = 6.38 \text{ m/s}$ . The difference in values is due to the rounding off errors in the numerical calculations of the mathematical quantities. If these two values are averaged the values are  $v_{1f} = -1.4 \text{ m/s}$  and  $v_{2f} = 6.6 \text{ m/s}$ ,  $p_f = 0.9 \text{ kg} \cdot \text{m/s}$ . Thus,  $p_i = p_f$ .

**IN THE MOVING FRAME:**

Make use of the Galilean velocity transformation equations.  $p'_i = m_1 v'_{1i} + m_2 v'_{2i}$ ; where  $v'_{1i} = v_{1i} - v' = 5 \text{ m/s} - (-2 \text{ m/s}) = 7 \text{ m/s}$ . Similarly,  $v'_{2i} = -1 \text{ m/s}$  and  $p'_i = 1.9 \text{ kg} \cdot \text{m/s}$ . To find  $p'_f$  use  $v'_{1f} = v_{1i} - v'$  and  $v'_{2f} = v_{2i} - v'$  because the prime system is now moving to the left. Using these results give  $p'_f = 1.9 \text{ kg} \cdot \text{m/s}$ .

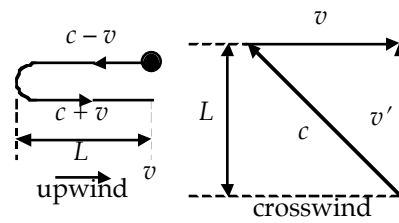
- 1-4 (a) In all cases one wants the speed of the plane relative to the ground. For the upwind and downwind legs, where  $v'$  in the figure is given by  $(c^2 - v^2)^{1/2}$

$$t_{u+d} = \frac{L}{c-v} + \frac{L}{c+v} = \frac{2L}{c} \left( \frac{1}{1-v^2/c^2} \right).$$

For the crosswind case, the plane's speed along  $L$  is  $v' = (c^2 - v^2)^{1/2}$

$$t_c = \frac{2L}{\sqrt{c^2 - v^2}} = \frac{2L}{c} \frac{1}{\sqrt{1 - (v/c)^2}}$$

$$t_{u+d} = \frac{2(100 \text{ mi})}{500 \text{ mi/h}} \left( \frac{1}{1 - (100)^2 / (500)^2} \right) = 0.4167 \text{ h}$$

$$t_c = \frac{2(100 \text{ mi})}{500 \text{ mi/h}} \left( \frac{1}{\sqrt{0.96}} \right) = 0.4082 \text{ h}$$


- (b)  $\Delta t = t_{u+d} - t_c = 0.0085 \text{ h} \approx 0.009 \text{ h}$  or  $0.510 \text{ min} \approx 0.5 \text{ min}$

- 1-5 This is a case of dilation.  $T = \gamma T'$  in this problem with the proper time  $T' = T_0$

$$T = \left[ 1 - \left( \frac{v}{c} \right)^2 \right]^{-1/2} T_0 \Rightarrow \frac{v}{c} = \left[ 1 - \left( \frac{T_0}{T} \right)^2 \right]^{1/2};$$

in this case  $T = 2T_0$ ,  $v = \left\{ 1 - \left[ \frac{T_0/2}{T_0} \right]^2 \right\}^{1/2} = \left[ 1 - \left( \frac{1}{4} \right) \right]^{1/2}$  therefore  $v = 0.866c$ .

- 1-6 This is a case of length contraction.  $L = \frac{L'}{\gamma}$  in this problem the proper length  $L' = L_0$ ,

$$L = \left[ 1 - \frac{v^2}{c^2} \right]^{-1/2} L_0 \Rightarrow v = c \left[ 1 - \left( \frac{L}{L_0} \right)^2 \right]^{1/2};$$

in this case  $L = \frac{L_0}{2}$ ,  $v = \left\{ 1 - \left[ \frac{L_0/2}{L_0} \right]^2 \right\}^{1/2} = \left[ 1 - \left( \frac{1}{4} \right) \right]^{1/2}$   
therefore  $v = 0.866c$ .

- 1-7 The problem is solved by using time dilation. This is also a case of  $v \ll c$  so the binomial expansion is used  $\Delta t = \gamma \Delta t' \approx \left[ 1 + \frac{v^2}{2c^2} \right] \Delta t'$ ,  $\Delta t - \Delta t' = \frac{v^2 \Delta t'}{2c^2}$ ;  $v = \left[ \frac{2c^2 (\Delta t - \Delta t')}{\Delta t'} \right]^{1/2}$ ;

$$\Delta t = (24 \text{ h/day})(3600 \text{ s/h}) = 86400 \text{ s}; \quad \Delta t' = \Delta t - 1 = 86399 \text{ s};$$

$$v = \left[ \frac{2(86\,400\text{ s} - 86\,399\text{ s})}{86\,399\text{ s}} \right]^{1/2} = 0.0048c = 1.44 \times 10^6\text{ m/s}.$$

1-8  $L = \frac{L'}{\gamma}$

$$\frac{1}{\gamma} = \frac{L}{L'} = \left[ 1 - \frac{v^2}{c^2} \right]^{1/2}$$

$$v = c \left[ 1 - \left( \frac{L}{L'} \right)^2 \right]^{1/2} = c \left[ 1 - \left( \frac{75}{100} \right)^2 \right]^{1/2} = 0.661c$$

1-10 (a)  $\tau = \gamma \tau'$  where  $\beta = \frac{v}{c}$  and

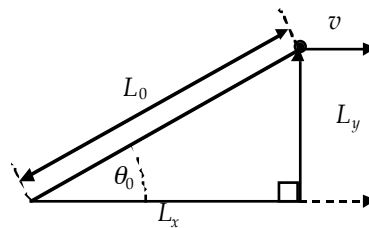
$$\gamma = (1 - \beta^2)^{-1/2} = \tau' \left( 1 - \frac{v^2}{c^2} \right)^{-1/2} = (2.6 \times 10^{-8}\text{ s}) \left[ 1 - (0.95)^2 \right]^{-1/2} = 8.33 \times 10^{-8}\text{ s}$$

(b)  $d = v\tau = (0.95)(3 \times 10^8)(8.33 \times 10^{-8}\text{ s}) = 24\text{ m}$

1-12 (a) 70 beats/min or  $\Delta t' = \frac{1}{70}\text{ min}$

(b)  $\Delta t = \gamma \Delta t' = \left[ 1 - (0.9)^2 \right]^{-1/2} \left( \frac{1}{70} \right)\text{ min} = 0.0328\text{ min/beat}$  or the number of beats per minute  $\approx 30.5 \approx 31$ .

1-14 (a) Only the x-component of  $L_0$  contracts.



$$L_{x'} = L_0 \cos \theta_0 \Rightarrow L_x = L_0 \cos \theta_0 / \gamma$$

$$L_{y'} = L_0 \sin \theta_0 \Rightarrow L_y = L_0 \sin \theta_0$$

$$L = \sqrt{L_x^2 + L_y^2} = \sqrt{L_0^2 \cos^2 \theta_0 / \gamma^2 + L_0^2 \sin^2 \theta_0}$$

$$= L_0 \sqrt{\cos^2 \theta_0 \left( 1 - \frac{v^2}{c^2} \right) + \sin^2 \theta_0} = L_0 \sqrt{1 - \frac{v^2}{c^2} \cos^2 \theta_0}$$

(b) As seen by the stationary observer,  $\tan \theta = \frac{L_y}{L_x} = \frac{L_0 \sin \theta_0}{L_0 \cos \theta_0 / \gamma} = \gamma \tan \theta_0$ .

- 1-16 For an observer approaching a light source,  $\lambda_{\text{ob}} = \left[ \frac{(1 - v/c)^{1/2}}{(1 + v/c)^{1/2}} \right] \lambda_{\text{source}}$ . Setting  $\beta = \frac{v}{c}$  and after some algebra we find,

$$\beta = \frac{\lambda_{\text{source}}^2 - \lambda_{\text{obs}}^2}{\lambda_{\text{source}}^2 + \lambda_{\text{obs}}^2} = \frac{(650 \text{ nm})^2 - (550 \text{ nm})^2}{(650 \text{ nm})^2 + (550 \text{ nm})^2} = 0.166$$

$$v = 0.166c = (4.98 \times 10^7 \text{ m/s})(2.237 \text{ mi/h})(\text{m/s})^{-1} = 1.11 \times 10^8 \text{ mi/h}.$$

- 1-17 (a) Galaxy A is approaching and as a consequence it exhibits blue shifted radiation.

From Example 1.6,  $\beta = \frac{v}{c} = \frac{\lambda_{\text{source}}^2 - \lambda_{\text{obs}}^2}{\lambda_{\text{source}}^2 + \lambda_{\text{obs}}^2}$  so that  $\beta = \frac{(550 \text{ nm})^2 - (450 \text{ nm})^2}{(550 \text{ nm})^2 + (450 \text{ nm})^2} = 0.198$ .

Galaxy A is approaching at  $v = 0.198c$ .

- (b) For a red shift, B is receding.  $\beta = \frac{v}{c} = \frac{\lambda_{\text{source}}^2 - \lambda_{\text{obs}}^2}{\lambda_{\text{source}}^2 + \lambda_{\text{obs}}^2}$  so that

$$\beta = \frac{(700 \text{ nm})^2 - (550 \text{ nm})^2}{(700 \text{ nm})^2 + (550 \text{ nm})^2} = 0.237. \text{ Galaxy B is receding at } v = 0.237c.$$

- 1-18 (a) Let  $f_c$  be the frequency as seen by the car. Thus,  $f_c = f_{\text{source}} \sqrt{\frac{c+v}{c-v}}$  and, if  $f$  is the frequency of the reflected wave,  $f = f_c \sqrt{\frac{c+v}{c-v}}$ . Combining these equations gives

$$f = f_{\text{source}} \frac{(c+v)}{(c-v)}.$$

- (b) Using the above result,  $f(c-v) = f_{\text{source}}(c+v)$ , which gives

$$(f - f_{\text{source}})c = (f + f_{\text{source}})v \approx 2f_{\text{source}}v.$$

The beat frequency is then  $f_{\text{beat}} = f - f_{\text{source}} = \frac{2f_{\text{source}}v}{c} = \frac{2v}{\lambda}$ .

- (c)  $f_{\text{beat}} = \frac{2(30.0 \text{ m/s})(10.0 \times 10^9 \text{ Hz})}{3.00 \times 10^8 \text{ m/s}} = \frac{2(30.0 \text{ m/s})}{0.0300 \text{ m}} = 2000 \text{ Hz} = 2.00 \text{ kHz}$

$$\lambda = \frac{c}{f_{\text{source}}} = \frac{3.00 \times 10^8 \text{ m/s}}{10.0 \times 10^9 \text{ Hz}} = 3.00 \text{ cm}$$

- (d)  $v = \frac{f_{\text{beat}}\lambda}{2}$  so,

$$\Delta v = \frac{\Delta f_{\text{beat}}\lambda}{2} = \frac{(5 \text{ Hz})(0.0300 \text{ m})}{2} = 0.0750 \text{ m/s} \approx 0.2 \text{ mi/h}$$